

→ Optomechanical Hamiltonian.



\hat{b} : mechanics
 \hat{a} : optics.

$$\hat{H} = \hbar \omega(\hat{x}) \hat{a}^\dagger \hat{a} + \hbar \Omega \hat{b}^\dagger \hat{b}$$

$$\hat{x} = x_{\text{ZPF}} (\hat{b}^\dagger + \hat{b}).$$

Typical cavity: $\omega(\hat{x}) = \omega_c \left[1 - \frac{\hat{x}}{L} + \mathcal{O}(\hat{x}^2) \right]$.

↳ 1st order \Rightarrow new term of the form $-\hat{F} \hat{x}$.
 with $\hat{F} = \frac{\hbar \omega_c}{L} \hat{a}^\dagger \hat{a}$.

$$\hat{H} = \hbar \omega_c \hat{a}^\dagger \hat{a} + \hbar \Omega \hat{b}^\dagger \hat{b} - \hbar \frac{\omega_c}{L} \hat{a}^\dagger \hat{a} \hat{x}$$

We define the coupling strength:

$$g_0 = \omega_c \frac{x_{\text{ZPF}}}{L}$$

More generally: $g_0 = -\frac{\partial \omega}{\partial x} x_{\text{ZPF}}$.

i.e. by how much does the zero point fluctuation changes the cavity resonance

→ What can we do with this interaction
 ↳ cooling / heating -

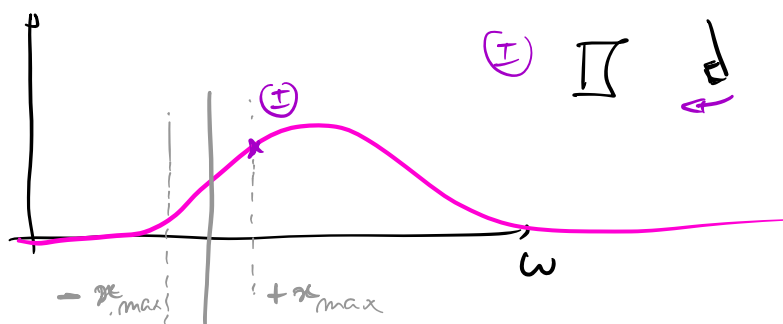
↳ first classical picture:
 How can we use the cavity to cool?

Interact: $m \rightarrow$ Force \propto # photon $\times x$
 \Rightarrow optical spring -

Mechanical spring constant:
 $D = D_0 - \frac{dF}{dx}$

$F \rightarrow kx$ form \Rightarrow conservative
 One needs to account for the lag
 in the cavity.

Q -factor of cavity \Rightarrow # photon
 Change at rate $1/\kappa$



 - longer cavity $\Rightarrow \omega_c \rightarrow$

- more photon when coming back
→ slow down
- fewer photon on way back
→ we don't re-accelerate.

↳ Cooling happens because
of lag-
the higher the lag, the better.

"Quantum picture":

In the rotating frame: $\hat{U} = e^{i\omega_L \hat{a}^\dagger \hat{a} t}$.

$$\hat{H} = \hbar \Delta \hat{a}^\dagger \hat{a} + \hbar \Omega \hat{b}^\dagger \hat{b} - \hbar g_0 \hat{a}^\dagger \hat{a} (\hat{b}^\dagger + \hat{b}).$$

$$\Delta = \omega_L - \omega_c. \quad \text{Hint.}$$

Linearisation: equivalent to applying displacement operator

Coherent amplitude: $\langle \hat{a} \rangle = \bar{\alpha}$

$$\hat{a} = \bar{\alpha} + \delta \hat{a}$$

$$H_{int} = -\hbar g_0 (\bar{\alpha} + \delta \hat{a})^\dagger (\bar{\alpha} + \delta \hat{a}) (\hat{b}^\dagger + \hat{b}).$$

Shift

$$= -\hbar g_0 |\bar{\alpha}|^2 (\hat{b}^\dagger + \hat{b}) - \hbar g_0 (\bar{\alpha}^* \hat{S} \hat{a} + \alpha \hat{S} \hat{a}^\dagger) (\hat{b}^\dagger + \hat{b}) - \hbar g_0 \hat{S} \hat{a}^\dagger \hat{S} \hat{a} (\hat{b}^\dagger + \hat{b}).$$

For large-ish field leading terms in $\bar{\alpha}$ are kept.

$-\hbar g_0 |\bar{\alpha}|^2 (\hat{b}^\dagger + \hat{b}) \rightarrow$ average shift of $\bar{\alpha}$ due radiation pressure.

Can be hidden by taking new reference for $\bar{\alpha}$.

It could be a problem if this shift is large compared to Kerr non-linearity of mechanics.

$$\hat{H}_{int} \approx -\hbar g_0 \sqrt{m_c} (\hat{S} \hat{a}^\dagger + \hat{S} \hat{a}) (\hat{b}^\dagger + \hat{b}).$$

$\hat{X} \hat{X}$ type interaction.

\hookrightarrow photon enhanced coupling $g \sqrt{m_c}$

Favourable regimes:

$\hookrightarrow g > \kappa \Rightarrow$ strong coupling
 \Rightarrow interaction faster than losses (in general $\Gamma_m \gg \kappa$)

$\hookrightarrow g_0 > \kappa \Rightarrow$ single-photon strong coupling.

interaction with 1 photon sufficiently strong \Rightarrow non-linear optomechanics.

\rightarrow Different RWA for different detunings:

$\Delta \approx -\Omega$, red side band:

$$H_{\text{int}} = -\hbar g (\hat{S} \hat{a}^\dagger \hat{b} + \hat{S} \hat{a} \hat{b}^\dagger).$$

\rightarrow cooling
 \rightarrow state transfer

) "beam-splitter" interaction

$\Delta \approx +\Omega$, blue side band:

$$H_{\text{int}} = -\hbar g (\hat{S} \hat{a}^\dagger \hat{b}^\dagger + \hat{S} \hat{a} \hat{b}).$$

\rightarrow amplification

\rightarrow entangler, squeezer

) "two-mode squeezer"

$\Delta \approx 0$, on resonance

$$H_{\text{int}} = -\hbar g (\hat{S} \hat{a}^\dagger + \hat{S} \hat{a}) (\hat{b}^\dagger + \hat{b})$$

\rightarrow position shift \Rightarrow phase shift

\rightarrow QND measurement of $\hat{S} \hat{a}^\dagger + \hat{S} \hat{a}$

→ Equation of motion.

↳ easiest way is through ϕ Langevin equations.

Both degree of freedom \hat{a}, \hat{b} are coupled to thermal bath (Fluctuations).
Treating the bath in Born-Markov approximation.

ϕ Langevin in short:

$$\hat{H} = \hat{H}_{\text{syst}} + \hat{H}_{\text{syst-bath}}.$$

Harmonic oscillators:

$$\hat{H}_{\text{syst}} = \frac{\hat{P}^2}{2m} + \hat{V}(\hat{q})$$

$$\hat{H}_{\text{syst-bath}} = \sum_j \left[\frac{\hat{P}_j^2}{2m_j} + \frac{k_j}{2} (\hat{q}_j - \hat{q})^2 \right]$$

coupling →

Heisenberg eq^o:

$$\dot{\hat{O}} = \frac{1}{i\hbar} [\hat{O}, \hat{H}_{\text{syst}}] + \frac{1}{i\hbar} [\hat{O}, \hat{H}_{\text{syst-bath}}]$$

easy |
$$= \frac{1}{i\hbar} [\hat{O}, \hat{H}_{\text{syst}}] - \frac{1}{2i\hbar} \sum_j k_j \left\{ [\hat{O}, \hat{q}] \hat{q}_j, \hat{q}_j - \hat{q} \right\}_+$$

$$\{A, B\}_+ = AB + BA.$$

Markovian of Langevin:

$$\dot{\hat{\mathcal{O}}} = \frac{1}{i\hbar} [\hat{\mathcal{O}}, \hat{H}_{\text{sys}}] + i\sqrt{2\gamma} [\hat{\mathcal{O}}, \hat{\rho}] \hat{P}_{\text{in}}(t) + \frac{1}{2i\gamma} \{ [\hat{\mathcal{O}}, \hat{Q}], \hat{\dot{Q}}(t) \}_+$$

$$\gamma = \frac{\mathcal{R}}{\gamma}.$$

For optomechanics:

$$\dot{\hat{a}} = -\frac{\kappa}{2} \hat{a} + i(\Delta - G \hat{x}) \hat{a} + \sqrt{\kappa_e} \hat{a}_{\text{in}} + \sqrt{\kappa_o} \hat{f}_{\text{in}}$$

$$\dot{\hat{b}} = -\frac{\Gamma}{2} \hat{b} - i\mathcal{R} \hat{b} + ig_0 \hat{a}^\dagger \hat{a} + \sqrt{\Gamma} \hat{b}_{\text{in}}.$$

$$G = \frac{g_0}{x_{\text{zpf}}}$$

$$\frac{1}{\kappa} = \frac{1}{\kappa_e} + \frac{1}{\kappa_o}.$$

With the noise correlations:

$$\langle \hat{a}_{\text{in}}(t) \hat{a}_{\text{in}}^\dagger(t') \rangle = \delta(t-t')$$

$$\langle \hat{a}_{\text{in}}^\dagger(t) \hat{a}_{\text{in}}(t') \rangle = 0$$

$$\langle \hat{b}_{\text{in}}(t) \hat{b}_{\text{in}}^\dagger(t') \rangle = (n_{\text{th}} + 1) \delta(t-t')$$

$$\langle \hat{b}_{\text{in}}^\dagger(t) \hat{b}_{\text{in}}(t') \rangle = n_{\text{th}} \delta(t-t').$$

) empty cavity!

empty cavity $\Leftrightarrow \frac{k_B T}{\hbar \omega_c} \approx 0.$

generally \mathcal{R} small $\Rightarrow \frac{k_B T}{\hbar \mathcal{R}} \gg 1. (n_{\text{th}} \gg 1)$

→ Classical version:

$$\langle \hat{a}(t) \rangle = \alpha(t)$$

$$\langle \hat{x}(t) \rangle = x(t).$$

$$\dot{\alpha} = -\kappa/2 \alpha + i(\Delta - G x) \alpha + \sqrt{\kappa_{ex}} \alpha_{in}$$

$$m_{eff} \ddot{x} = -m_{eff} \Omega^2 x - m_{eff} \Gamma \dot{x} + \hbar G |\alpha|^2$$



Full Quantum Langevin, in practice not solvable. To access the q fluctuation, one can apply the same linearisation as for μ : $\hat{a} = \alpha + \delta \hat{a}$. $\alpha = \sqrt{n_c}$.

→ Linearised Classical Langevin:

$$\delta \dot{\alpha} = (i\Delta - \frac{\kappa}{2}) \delta \alpha + i G \bar{\alpha} x$$

$$m_{eff} \ddot{x} = -m_{eff} \Omega^2 x - m_{eff} \Gamma \dot{x} + \hbar G (\bar{\alpha}^* \delta \alpha + \bar{\alpha} \delta \alpha^*)$$

Fourier transform:

$$-i\omega \delta \dot{\alpha} = (i\Delta - \frac{\kappa}{2}) \delta \alpha + i G \bar{\alpha} x$$

$$-m_{eff} \omega^2 \ddot{x} = -m_{eff} \Omega^2 x + i\omega m_{eff} \Gamma \dot{x} + \hbar G (\bar{\alpha}^* \delta \alpha + \bar{\alpha} \delta \alpha^*)$$

$G=0 \Rightarrow$ mechanical susceptibility,

$$\bullet \chi_{xx}^{-1}(\omega) = m_{eff} [(\Omega^2 - \omega^2) - i\Gamma\omega].$$

Coupling to the cavity ($G \neq 0$)

$$\Rightarrow \chi_{nn}^{-1} = \chi_{nn,0}^{-1}(\omega) + \Sigma(\omega)$$

Self energy.

the mechanical susceptibility is changed by the radiation pressure. K is the cavity lag.

Solving for $\bar{\alpha}$ and $\delta\alpha$ we get:

$$\Sigma(\omega) = 2 m_{\text{eff}} \Omega g^2 \left\{ \frac{1}{(\Delta + \omega) + iK/2} + \frac{1}{(\Delta - \omega) - iK/2} \right\}.$$

The real and imaginary part have \neq roles.

$\hookrightarrow \chi_{nn}(\omega) \xrightarrow{\text{real}} \Rightarrow \text{mech. freq.}$
 $\searrow \text{imag} \Rightarrow \text{mech damping}$

$$\Sigma(\omega) = m_{\text{eff}} \omega \left[2\delta\Omega(\omega) - i\Gamma_{\text{opt}}(\omega) \right].$$

$$\hookrightarrow \chi_{nn}^{-1}(\omega) = m_{\text{eff}} \left[\Omega^2 + 2\omega\delta\Omega - \omega^2 - i\omega[\Gamma + \Gamma_{\text{opt}}(\omega)] \right]$$

$$\left[\begin{aligned} \delta\Omega &= g^2 \frac{\Omega}{\omega} \left[\frac{\Delta + \omega}{(\Delta + \omega)^2 + K^2/4} + \frac{\Delta - \omega}{(\Delta - \omega)^2 + K^2/4} \right] \\ \Gamma_{\text{opt}} &= g^2 \frac{\Omega}{\omega} \left[\frac{K}{(\Delta + \omega)^2 + K^2/4} - \frac{K}{(\Delta - \omega)^2 + K^2/4} \right] \end{aligned} \right.$$

→ Optical spring : $\delta\Omega$ @ $\omega = \Omega$.

$$\delta\Omega = g_0^2 \bar{m}_c \left[\frac{\Delta + \Omega}{(\Delta + \Omega)^2 + \kappa^2/4} + \frac{\Delta - \Omega}{(\Delta - \Omega)^2 + \kappa^2/4} \right].$$

→ Optical damping :

$$\Gamma_{\text{opt}} = g_0^2 \bar{m}_c \left[\frac{\kappa}{(\Delta + \Omega)^2 + \kappa^2/4} - \frac{\kappa}{(\Delta - \Omega)^2 + \kappa^2/4} \right].$$

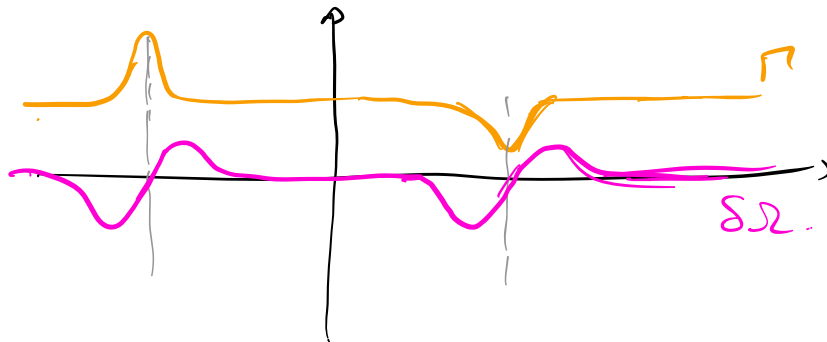
$$\Gamma_{\text{eff}} = \Gamma + \Gamma_{\text{opt}}.$$

Drawing the different ones.

Resolved sideband

Unresolved ———.

$\kappa \ll \Omega$



Why $\Gamma_{\text{eff}} \gg \Gamma \Rightarrow$ cooling.
Extra damping from a "cold" cavity.

$$\Gamma_{\text{opt}} |_{\kappa \ll \Omega} = \frac{4 g_0^2}{\kappa} \bar{m}_c.$$

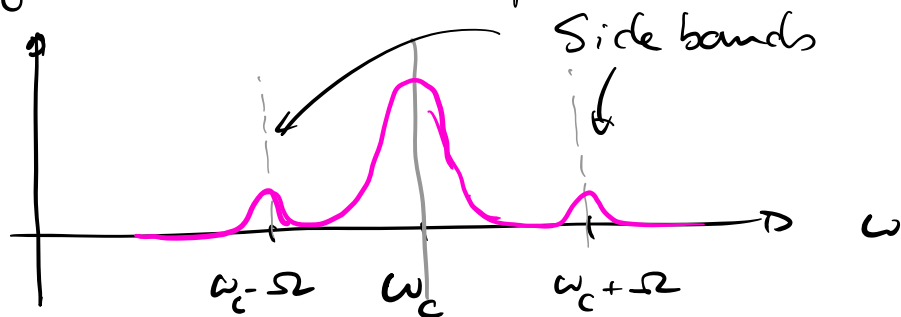
cooperativity : $\mathcal{G}_0 = \frac{4 g_0^2}{\kappa \Gamma}$
 $\mathcal{G} = \Gamma_{\text{opt}} / \Gamma$

↳ Back with the hands:

cavity lag \Rightarrow dissipative force.

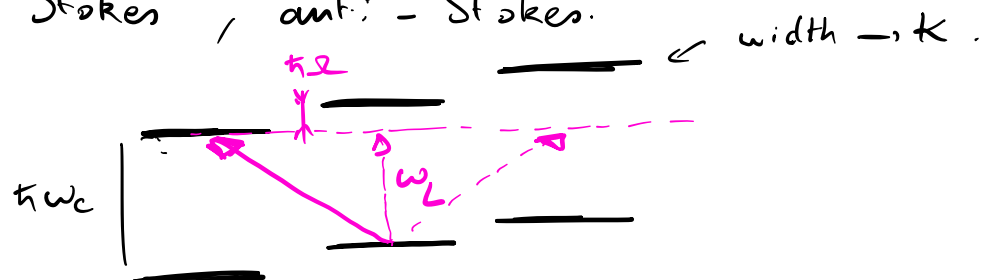
Why do we want $\hbar \ll \hbar \Omega$ (resolved sidebands)

Cavity + mechanics response:



↳ we can understand this as

Stokes, anti-Stokes.



↳ Cooling limits:

"naive"

picture:

$$T_{\text{final}} = T_{\text{init}} \frac{\Gamma}{\Gamma + \Gamma_{\text{opt}}}$$

$$\Gamma_{\text{opt}} \xrightarrow{\hbar \Omega \rightarrow \infty} \infty \Rightarrow T_{\text{final}} \rightarrow 0.$$

We are missing shot noise again

two processes:

$$\begin{aligned}\Gamma_{n \rightarrow n-1} &= n A^- \\ \Gamma_{n \rightarrow n+1} &= (n+1) A^+\end{aligned}$$

Γ_{opt} simply: $\Gamma_{\text{opt}} = A^- - A^+$.

Evolution of phonon #: \bar{n}_p .

$$\begin{aligned}\dot{\bar{n}}_p &= (\bar{n}_p + 1)(A^+ + A_{\text{th}}^+) - \bar{n}_p(A^- + A_{\text{th}}^-) \\ A_{\text{th}}^+ &= \bar{n}_{\text{th}} \Gamma \quad \text{if } \bar{n}_{\text{th}} = 0 \text{ can't gain} \\ A_{\text{th}}^- &= (\bar{n}_{\text{th}} + 1) \Gamma\end{aligned}$$

@ equilibrium: $\dot{\bar{n}}_p = 0$.

$$\bar{n}_p = \frac{A^+ + \bar{n}_{\text{th}} \Gamma}{\Gamma_{\text{opt}} + \Gamma} < \bar{n}_{\text{min}} = \frac{A^+}{\underbrace{A^- - A^+}_{\Gamma_{\text{opt}}}}$$

We can calculate A^\pm with the Fermi Golden rule:

Forcing term: radiation pressure from cavity field.

$$\hat{H}_{\text{int}} = -\hat{a} \hat{F} \quad \text{with} \quad \hat{F} = \hbar G \hat{a}^\dagger \hat{a}.$$

recalling: $S_{FF}(\omega) = \int dt e^{i\omega t} \langle \hat{F}(t) \hat{F}(0) \rangle$

$$A^{\pm} = \frac{\alpha_{\text{ZPF}}^2}{\hbar^2} S_{\text{FF}}(\omega = \mp \Omega) = g_0^2 S_{\text{NN}}(\omega = \mp \Omega)$$

the normalised spectrum is simply:
 $S_{\text{NN}}(\omega) = \int dt e^{i\omega t} \langle \hat{a}^\dagger \hat{a}(t) \hat{a}^\dagger \hat{a}(0) \rangle.$

$$= \bar{n}_c \frac{\kappa}{\kappa/4 + (\Delta + \omega)^2}$$

$$\text{Finally: } \bar{n}_{\text{min}} = \left(\frac{A^-}{A^+} - 1 \right)^{-1} = \left(\frac{(\kappa/2)^2 + (\Delta - \Omega)^2}{(\kappa/2)^2 + (\Delta + \Omega)^2} - 1 \right)^{-1}$$

$$\rightarrow \kappa \ll \Omega \Rightarrow \bar{n}_{\text{min}} = \left(\frac{\kappa}{4\Omega} \right)^2 < 1$$

$$\rightarrow \kappa \gg \Omega \Rightarrow \bar{n}_{\text{min}} = \frac{\kappa}{4\Omega} \gg 1$$

Careful this value is for $\Delta = -\Omega$
 not optimum in unresolved regime

→ Simon entanglement

$$\rightarrow |\psi_{\text{EPR}}\rangle \propto |00\rangle_{AB}$$

$$+ \sqrt{p_b} (|H\rangle|01\rangle_{AB} + |V\rangle|10\rangle_{AB})$$

$$\rightarrow |\psi_{\text{in}}\rangle \propto |0\rangle + \alpha \left(\cos \frac{\theta_{\text{in}}}{2} |H\rangle + \sin \frac{\theta_{\text{in}}}{2} |V\rangle \right) + \mathcal{O}(\alpha^2)$$

$$\rightarrow |\psi_{\text{out}}\rangle = \cos \frac{\theta_{\text{in}}}{2} |10\rangle_{AB} \pm e^{i\phi_m} \sin \frac{\theta_{\text{in}}}{2} |01\rangle_{AB}$$